



EUROPEAN ARC  
ALMA Regional Centre || Allegro



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# ANALYSIS TOOLS



Slides taken from the 15th Synthesis  
Imaging Workshop by Ylva Pihlstr.m

[science.nrao.edu/science/meetings/2016/  
15th-synthesis-imaging-workshop/  
SIW2016\\_Pihlstrom.pdf](http://science.nrao.edu/science/meetings/2016/15th-synthesis-imaging-workshop/SIW2016_Pihlstrom.pdf)

# Visualizing

- Imaging will create a spectral line *cube*, which is 3-dimensional: RA, Dec and Velocity.
- With the cube, we usually visualize the information by making 1-D or 2-D projections:
  - Line profiles (1-D slices along velocity axis)
  - Channel maps (2-D slices along velocity axis)
    - 'Movies' can be formed from the channel maps
  - Moment maps (integration along the velocity axis)
  - Position-velocity plots (slices along spatial dimension)
- 3-D rendering programs also exist

# Visualizing 3-D

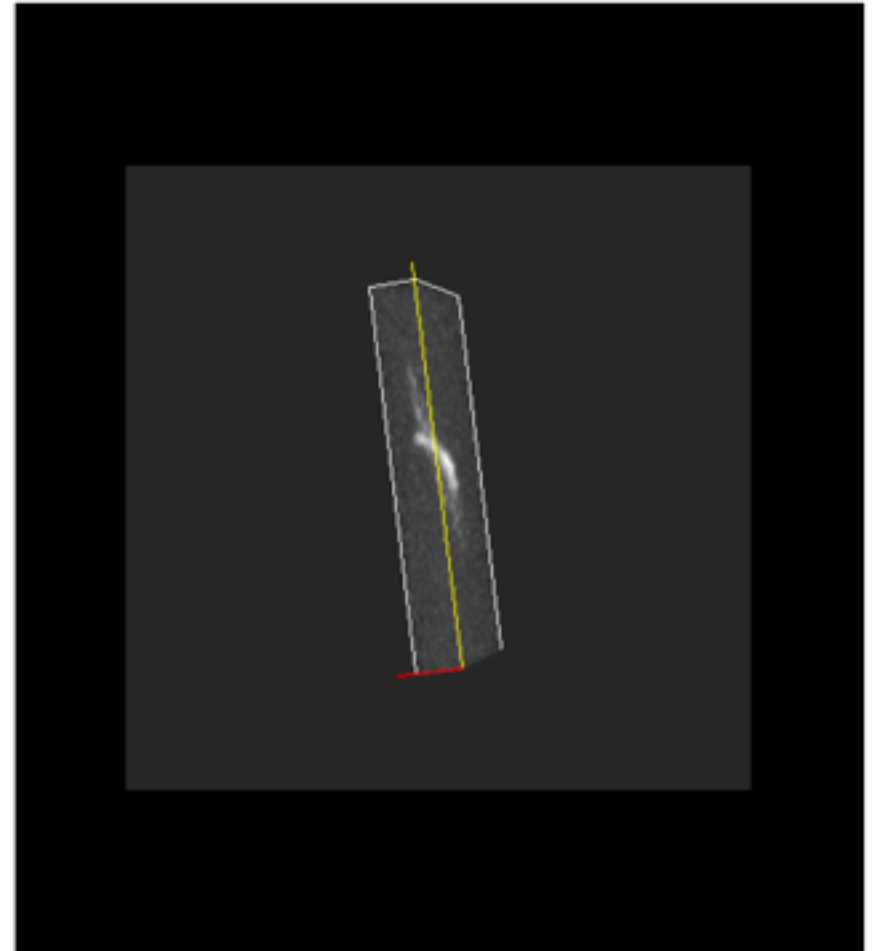
Some software allows 3-D visualization of cubes:

SAOimage DS9:

<http://ds9.si.edu/site/Home.html>

Karma package:

<http://www.atnf.csiro.au/computing/software/karma/>

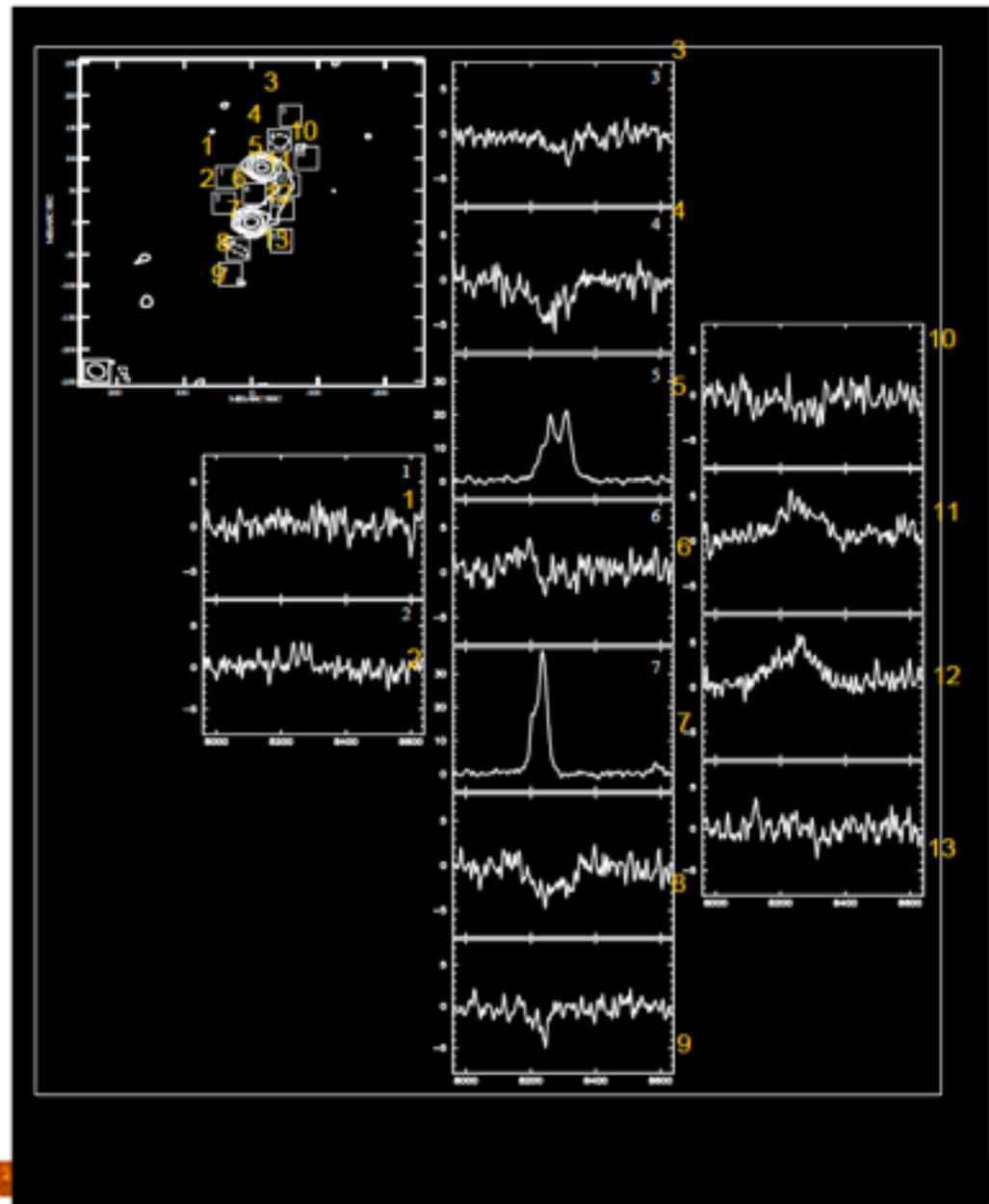


Karma package; L. Matthews

# I-D line profiles

- Line profiles show changes in line shape, width and depth as a function of position.
- Can give information of relative position of features (absorption in front)
- Velocity width

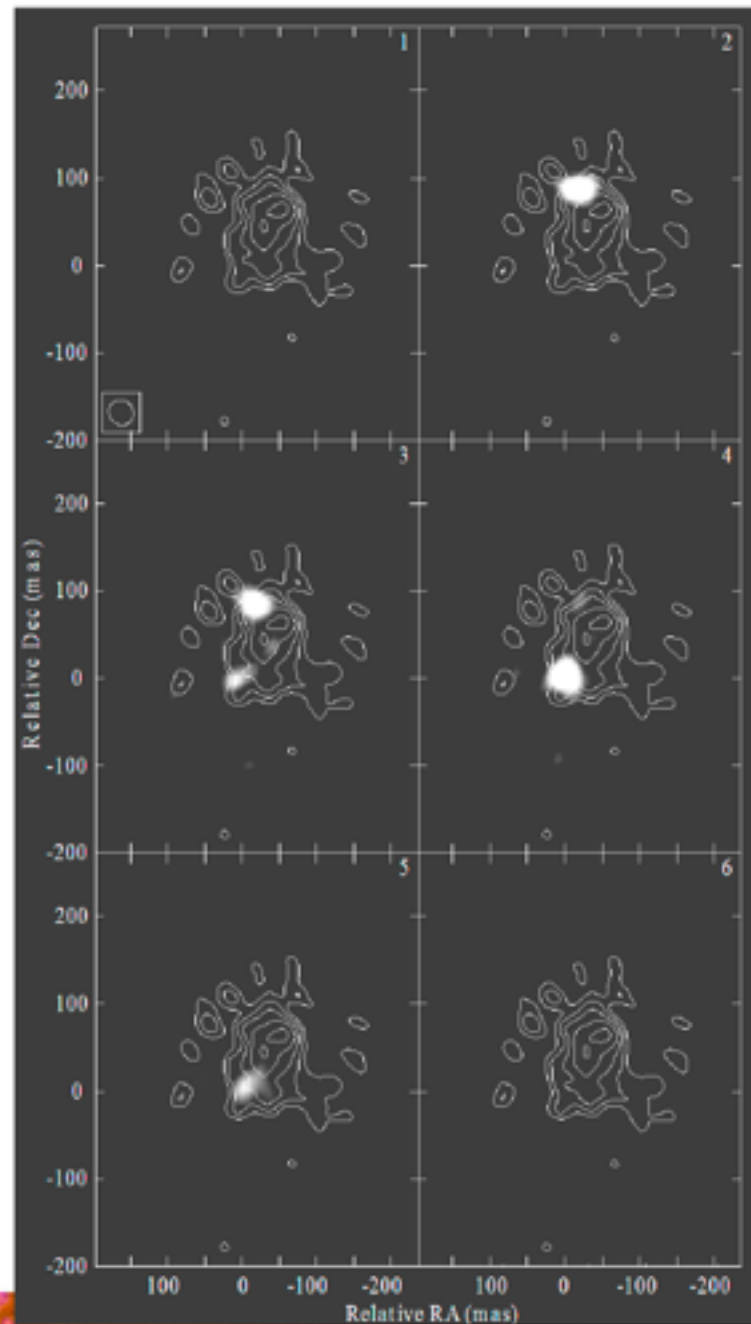
EVN+MERLIN 1667 MHz  
OH maser emission and  
absorption spectra in a  
luminous infrared galaxy  
(IIIZw35).



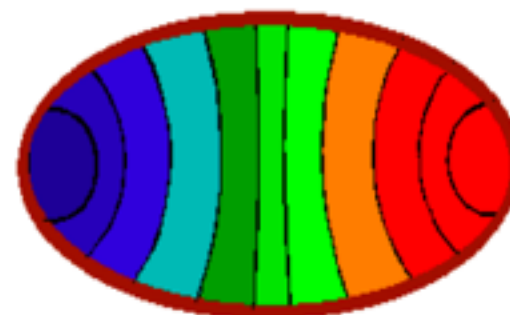
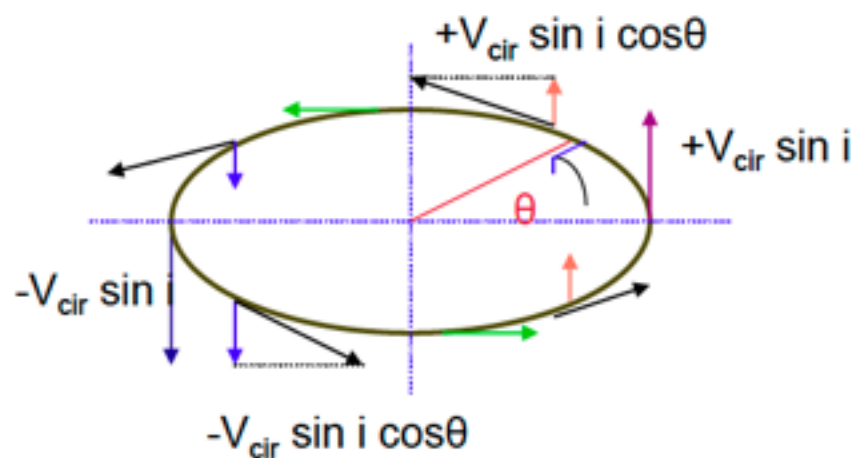
## 2-D channel maps

- Channel maps show how the spatial distribution of the line feature changes with frequency/velocity.
- Information about kinematics.

Contours continuum  
emission, grey scale  
1667 MHz OH line  
emission in III Zw35.

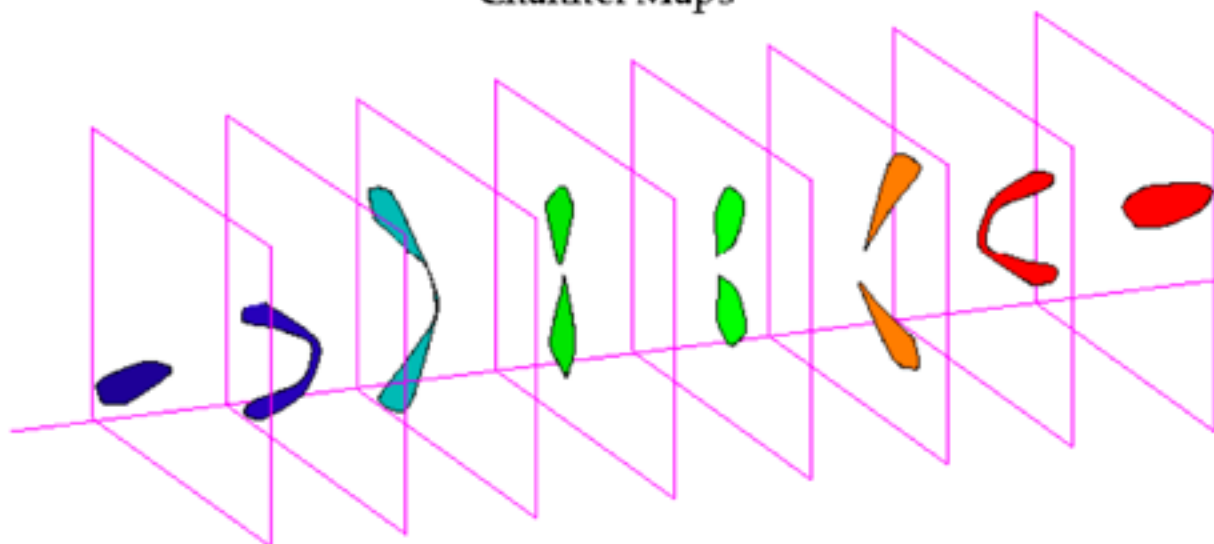


# Rotating disk model



Mean Velocity Field

Channel Maps

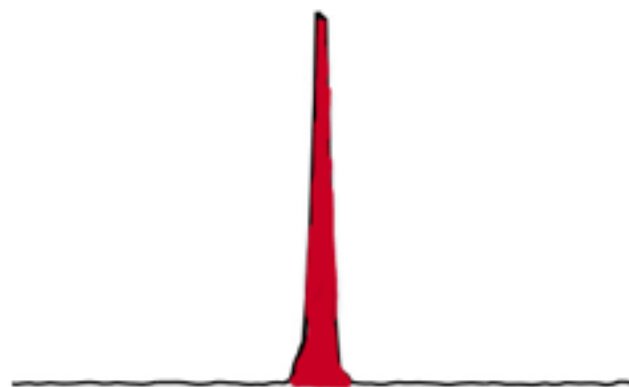




# Moment analysis

- You might want to derive parameters such as integrated line intensity, centroid velocity of components and line width - all as functions of positions. Estimate using the *moments* of the line profile:
- Integrated intensity

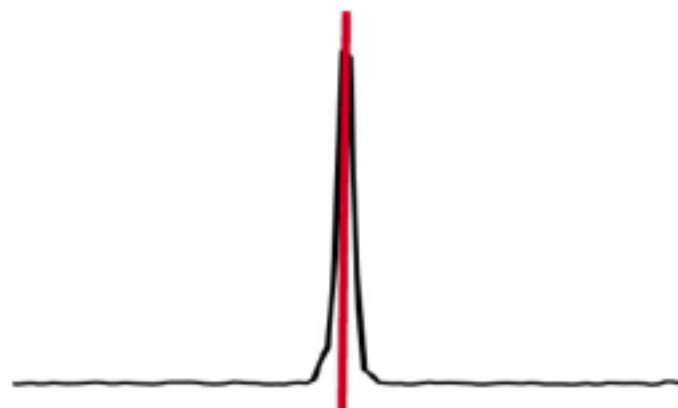
$$\text{Moment } 0 = \int S_v dv$$





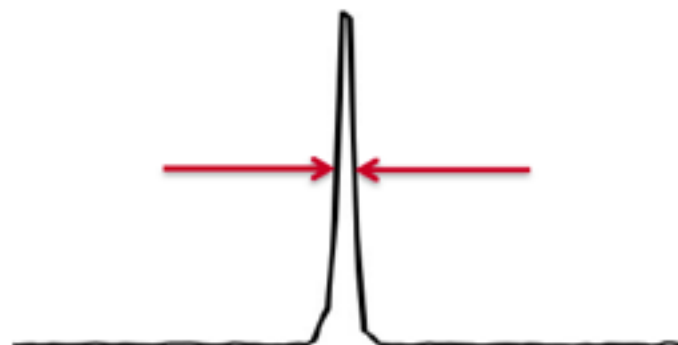
- Intensity weighted velocity

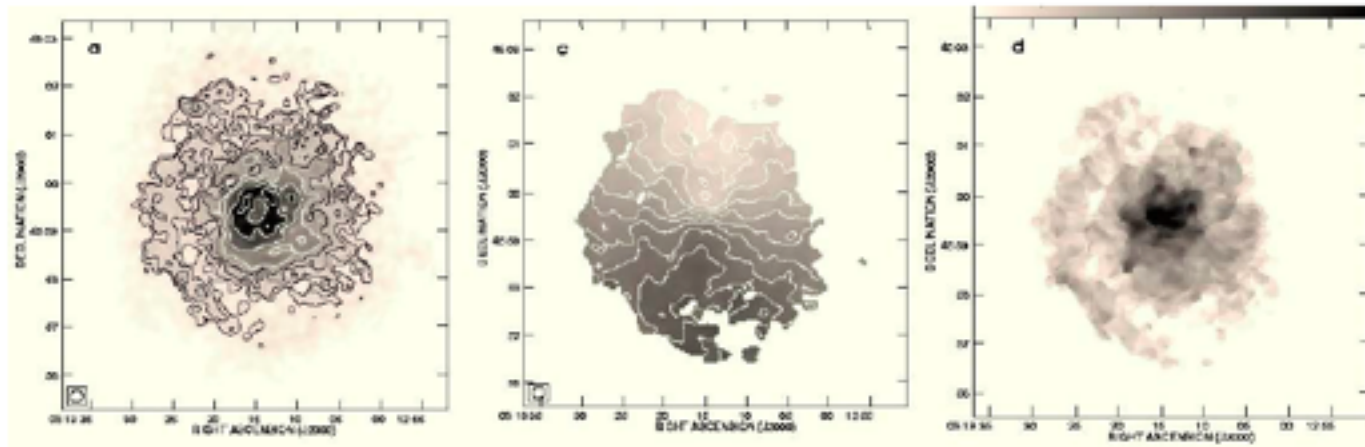
$$\text{Moment 1} = \langle V \rangle = \frac{\int S_v v dv}{\int S_v dv}$$



- Intensity weighted velocity dispersion

$$\text{Moment 2} = \langle V^2 \rangle^{1/2} = \sqrt{\frac{\int S_v (v - \langle V \rangle)^2 dv}{\int S_v dv}}$$





Moment 0  
Total intensity

Moment 1  
Velocity field

Moment 2  
Velocity dispersion

- Moments sensitive to noise so clipping is required
- Higher order moments depend on lower ones so progressively noisier.

**XMOM, MOMNT/immoments**

# Image Analysis of TWHydra

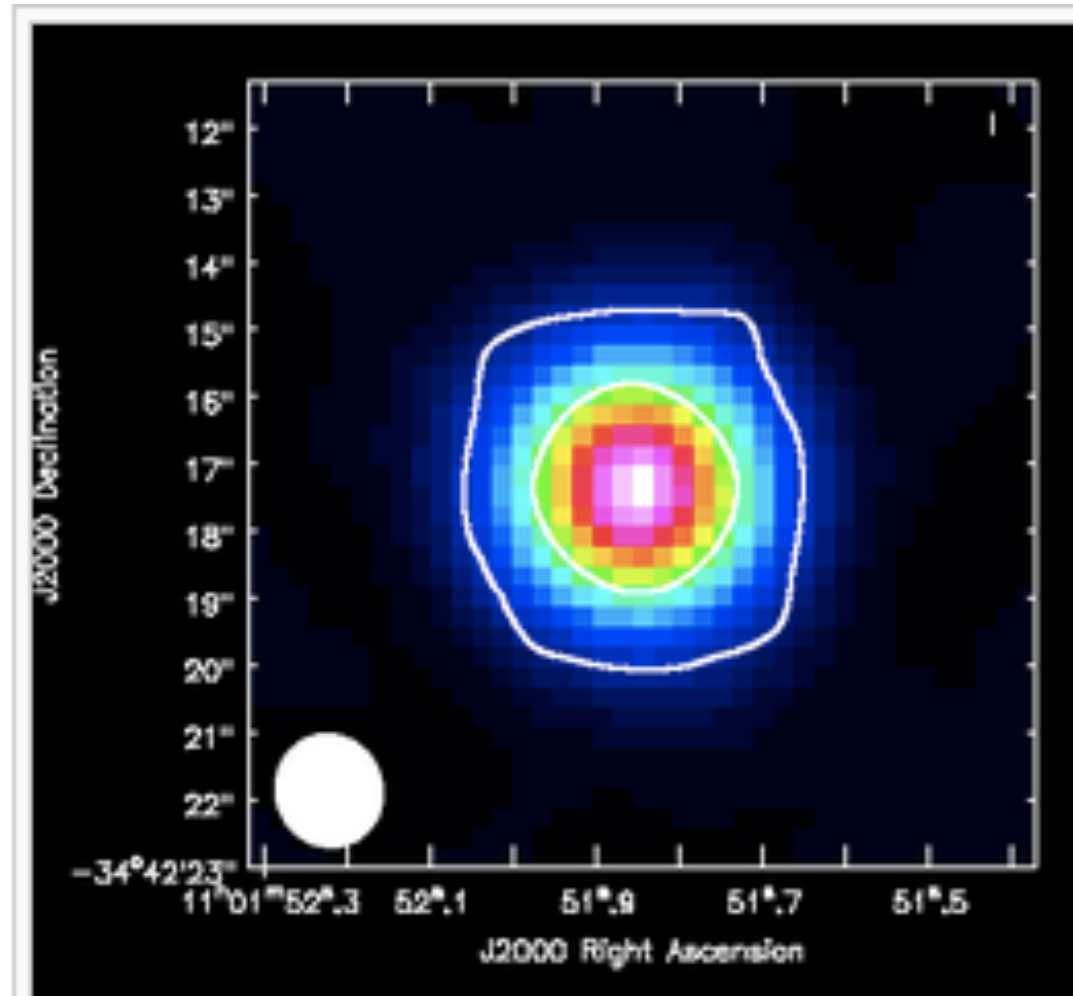


- Using the CASA guides: [https://casaguides.nrao.edu/index.php/First\\_Look\\_at\\_Image\\_Analysis](https://casaguides.nrao.edu/index.php/First_Look_at_Image_Analysis)
- Using the CASA guides: [https://casaguides.nrao.edu/index.php/TWHydraBand7\\_Imaging\\_4.5#Image\\_Analysis](https://casaguides.nrao.edu/index.php/TWHydraBand7_Imaging_4.5#Image_Analysis)

# Moment 0



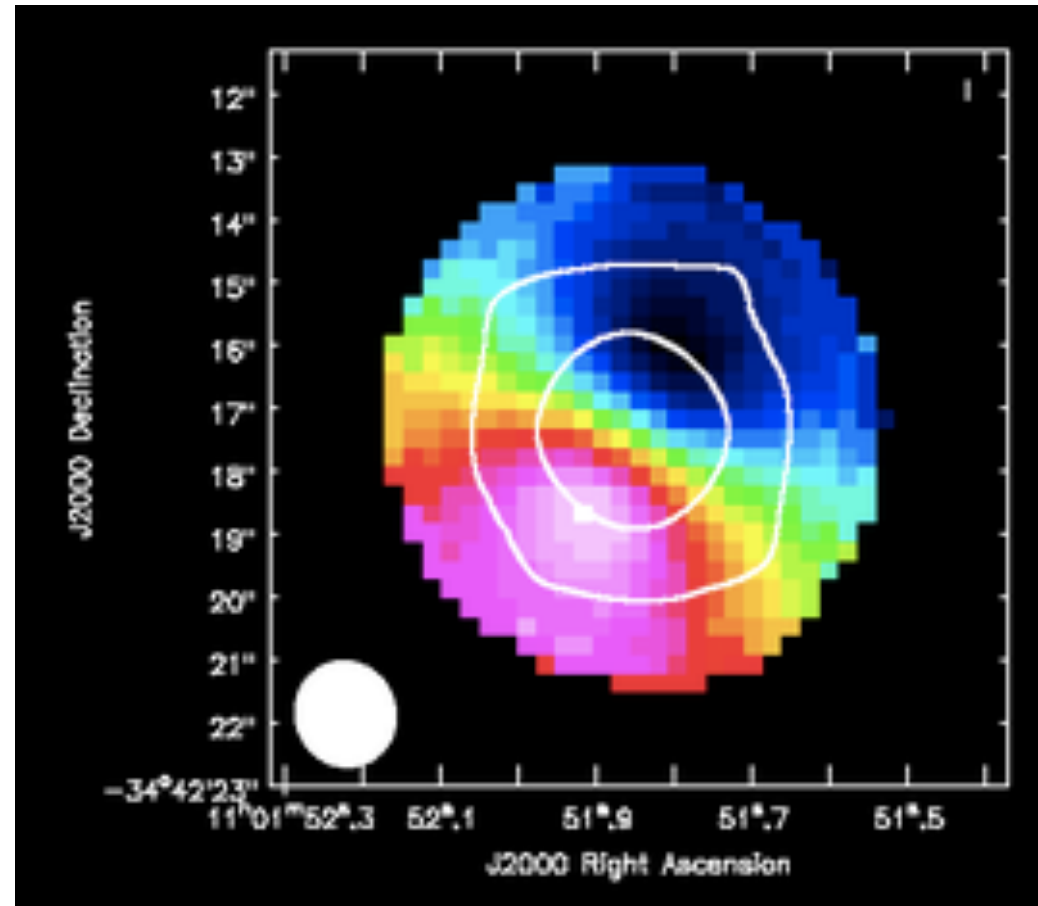
```
os.system("rm -rf  
sis14_twhya_n2hp.mom0")  
  
immoments("sis14_twhya_n2hp.im  
age",  
outfile="sis14_twhya_n2hp.mom0",  
includepix=[20e-3,100],  
chans="4~12", moments=0)
```



# Moment 1



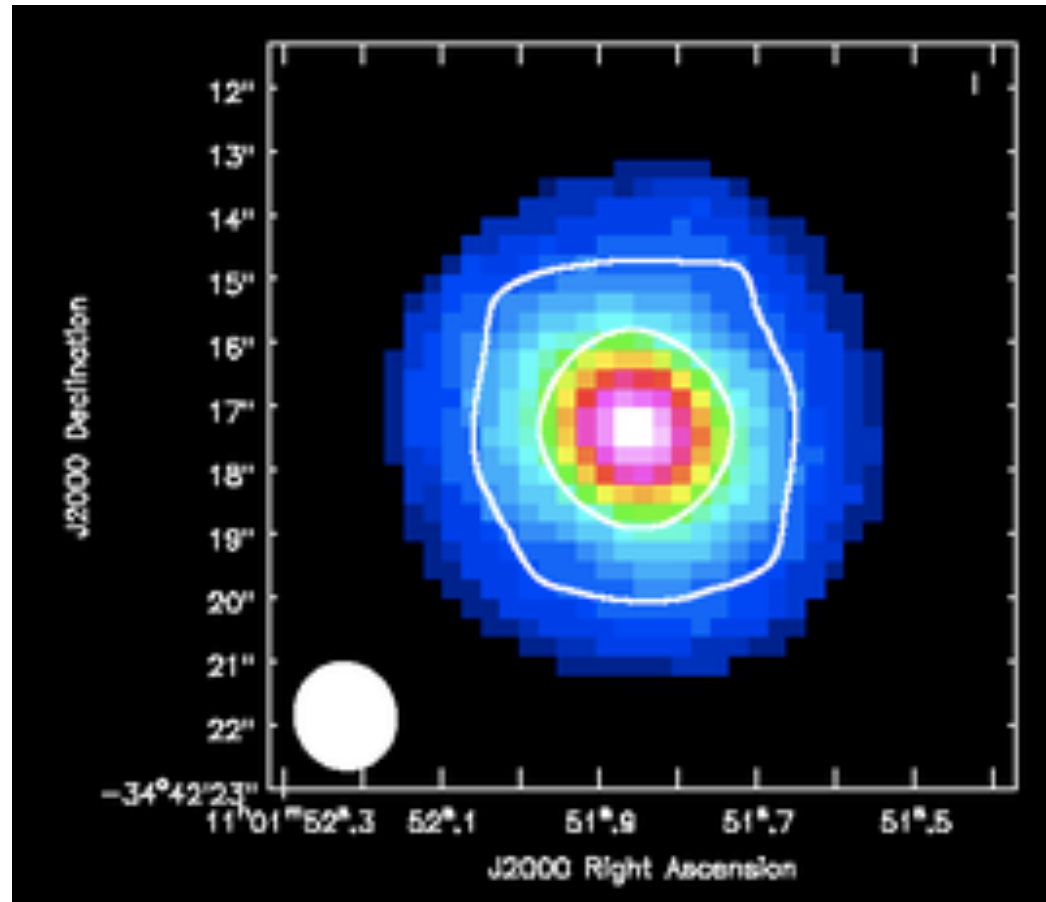
```
os.system("rm -rf  
sis14_twhya_n2hp.mom1")  
  
immoments("sis14_twhya_n2hp.im  
age",  
outfile="sis14_twhya_n2hp.mom1",  
includepix=[40e-3,100],  
chans="4~12", moments=1)
```



# Moment 2



```
os.system("rm -rf  
sis14_twhya_n2hp.mom1")  
immoments("sis14_twhya_n2hp.im  
age",  
outfile="sis14_twhya_n2hp.mom1",  
includepix=[40e-3,100],  
chans="4~12", moments=2)
```



# Links



<https://casa.nrao.edu/docs/cookbook/>

Image analysis chapter:

[https://casa.nrao.edu/docs/cookbook/  
casa\\_cookbook007.html](https://casa.nrao.edu/docs/cookbook/casa_cookbook007.html)

CASA toolkit:

<https://casa.nrao.edu/docs/CasaRef/image.moments.html>