

ALMA DATA

FROM THE ARCHIVE TO CALIBRATED VISIBILITIES

ALMA Data Reduction Training Day

Aida Ahmadi

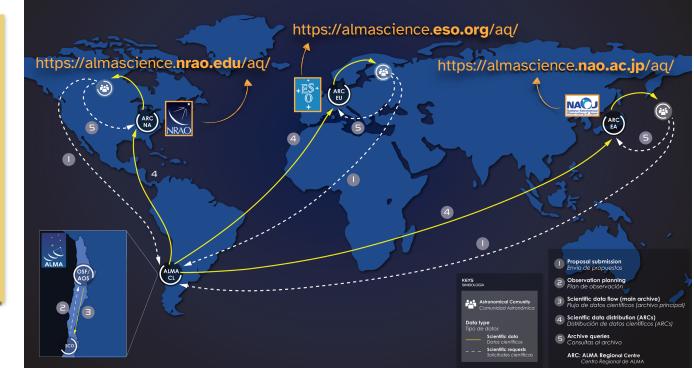
ALMA Local Expertise Group (Allegro)



Leiden Observatory November 27, 2023

The ALMA Science Archive

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ASA includes:

- science verification
- regular projects
- DDT projects
- large programs
- projects that use ALMA as part of VLBI
- calibrators

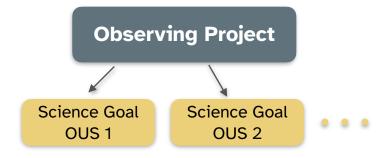
https://almascience.eso.org/aq/

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ALMA Data Reduction Training Day

November 27, 2023

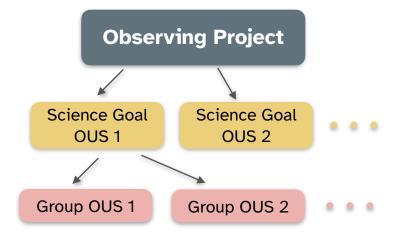
- **OUS:** Observing Unit Set smallest unit
- Science Goal: Defined by the PI in the observing tool (OT) at the proposal stage





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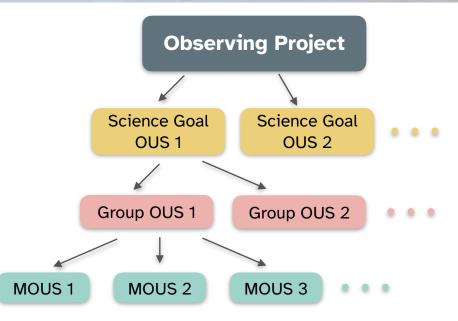
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- **Group OUS:** A group can contain several configurations to be combined in data processing





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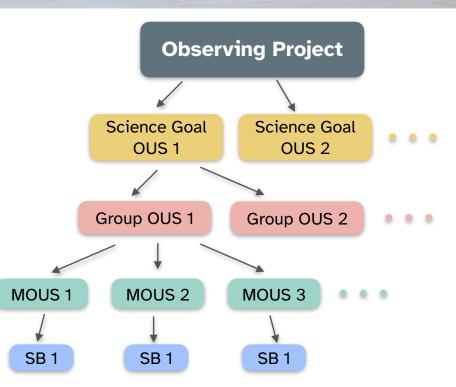
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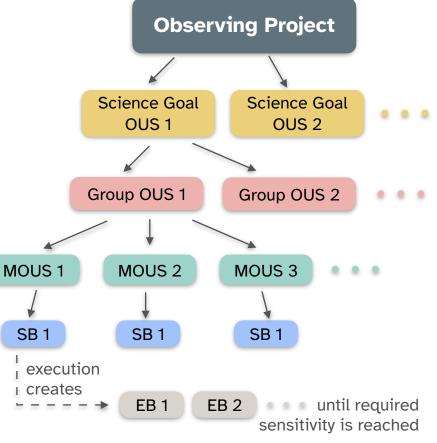
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- Scheduling Block (SB): smallest entity used for observing. It is a plan for a complete set of calibration and science target observing scans





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- **Group OUS:** A group can contain several configurations to be combined in data processing
- Member OUS (MOUS): all executions of an SB
- Scheduling Block (SB): smallest entity used for observing. It is a plan for a complete set of calibration and science target observing scans
- Execution Block (EB): each repetition of a scheduling block (SB)



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Quality Assurance (QA)

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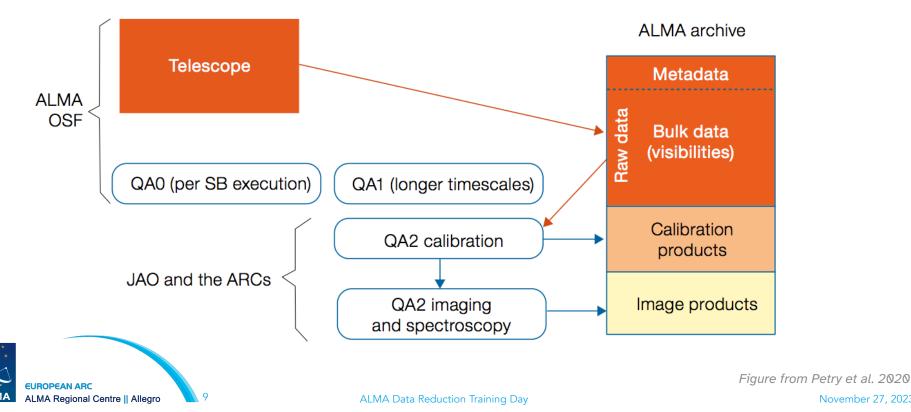
QA consists of 3 (+1) steps:

- QAO(+): performed at the telescope shorty after execution of a SB -> check the correct setup of antennas & receivers, stability of atmosphere, verifying that the flux calibrator used has a recent flux measurement
- **QA1:** longer-term monitoring of observatory parameters
- **QA2:** offline calibration and imaging on MOUSs to confirm the science goal was met
 - If the requested sensitivity & angular resolution achieved -> data delivered
 - If not (<10% of cases): re-observe SB & new QA2 process until requests are met
- QA3: (optional) triggered if errors are discovered by the PI or ALMA staff after data delivery



ALMA Data Flow

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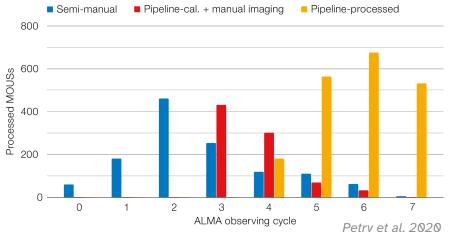
Pipeline processing

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- Earlier cycles: QA2 processing exclusively done semi-manually by analysts in CASA and the Calibration Script Generator
- Since then: fully automated pipeline -> distributed with CASA releases
- Weblogs:

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- The pipeline creates a set of \bigcirc diagnostic plots and tables
- reviewed manually to judge \bigcirc whether the pipeline run was successful, and the observing parameters were met



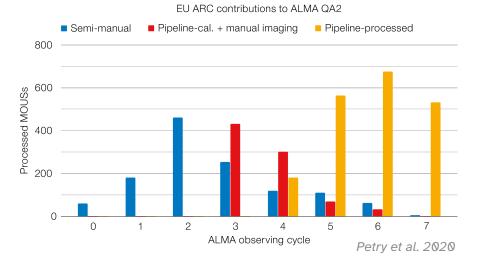


EU ARC contributions to ALMA QA2

Pipeline processing

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- <10% of cases require semi-manual processing by analysts
- 90% of the deliveries done within 1 month after the observation
 - Median of 2 weeks
- calibrated visibilities & single-dish data are not stored in the Archive & are not part of the data delivery





What exactly is on the archive?

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- **README:** description of the main actions, categories, and files
- product: FITS images produced by the pipeline process (includes calibrator & target)
- auxiliary: calibration tables, logs, data quality plots, scripts
- qa: Quality Assurance files and plots -> weblog!
- raw data: ALMA Science Data Model (ASDM) files, one per EB of QA2 Pass/Semi-pass
- raw (semipass): ASDM files of QA0-Semipassed data
- external: usually imaging products

Depending on the cycle, the delivered products may vary... Check for each cycle: <u>https://almascience.org/processing/ga2-data-products</u>



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Data on the ALMA Science Archive

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- The calibrated datasets on the archive are science-ready
- Users are strongly encouraged to re-image the calibrated datasets manually
- Not all image products are on the archive, but typically
 - continuum-subtracted image cubes at the requested resolution
 - continuum image either combining all SPWs or all line-free channels
- ARI-L project: aims to increase the legacy value of the ASA by reprocessing Cycle 2-4 data to bring them close to the level of more recent cycles (Massardi et al. 2019) -> <u>https://sites.google.com/inaf.it/ari-l/project</u>



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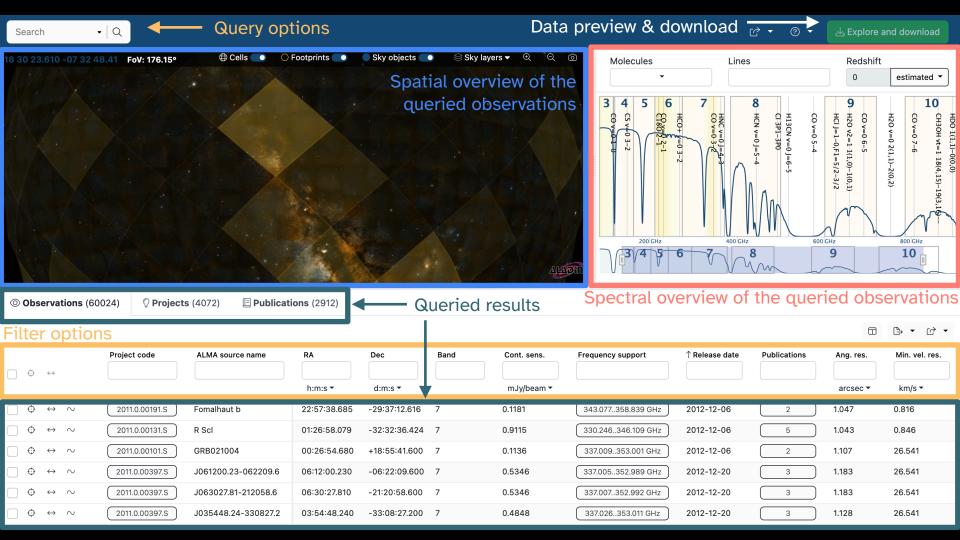
Large Programs: Enhanced data products

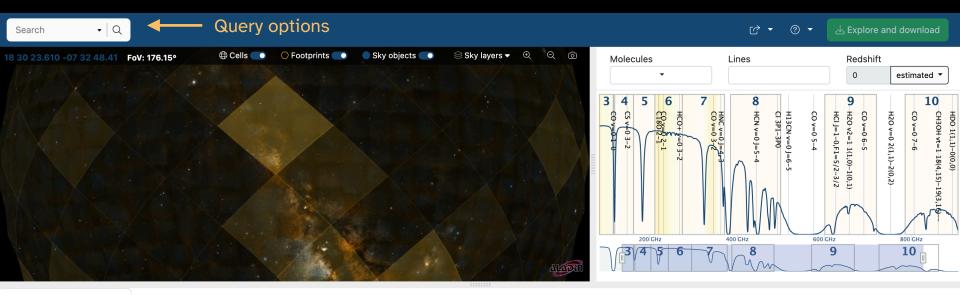
https://almascience.eso.org/alma-data/lp



Let's go to the archive... https://almascience.eso.org/aq/







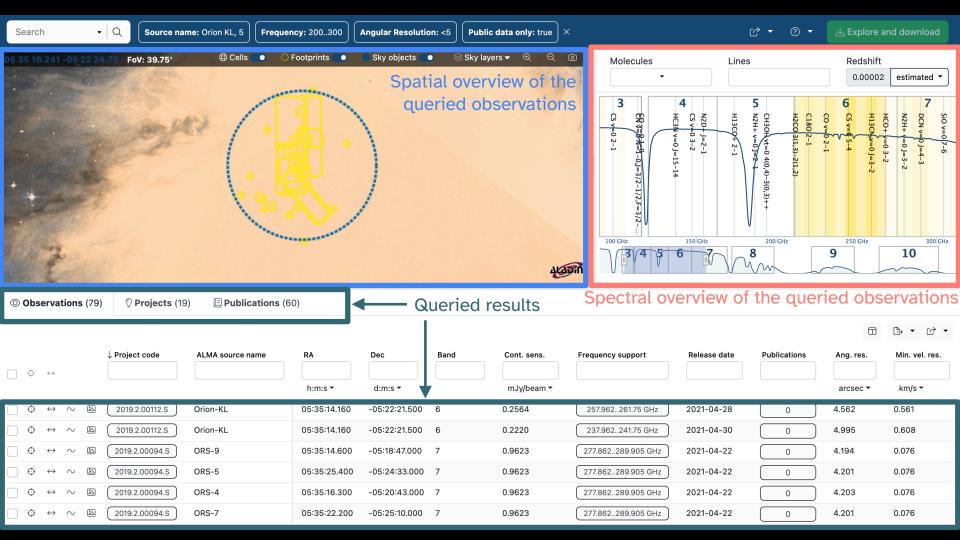
Observations (60024)

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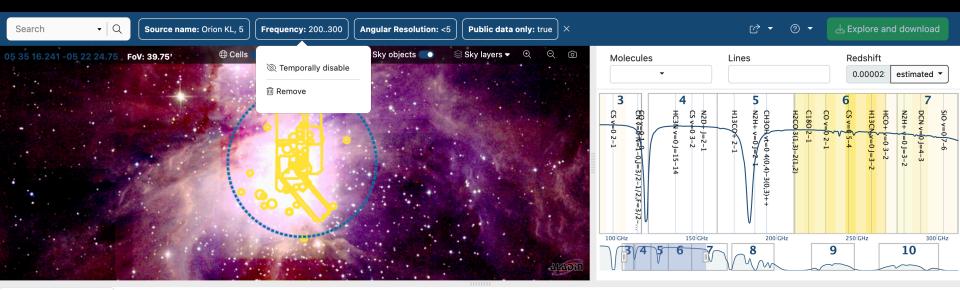
$\Box \Leftrightarrow \leftrightarrow$	Project code	ALMA source name	RA	Dec	Band	Cont. sens.	Frequency support	↑ Release date	Publications	Ang. res.	Min. vel. res.
			h:m:s ▼	d:m:s ▼		mJy/beam ▼				arcsec 🔻	km/s ▼
$\Box \Leftrightarrow \leftrightarrow \sim$	2011.0.00191.S	Fomalhaut b	22:57:38.685	-29:37:12.616	7	0.1181	343.077358.839 GHz	2012-12-06	2	1.047	0.816
$\Box \oplus \leftrightarrow \sim$	2011.0.00131.S	R Scl	01:26:58.079	-32:32:36.424	7	0.9115	330.246346.109 GHz	2012-12-06	5	1.043	0.846
$\Box \oplus \leftrightarrow \sim$	2011.0.00101.S	GRB021004	00:26:54.680	+18:55:41.600	7	0.1136	337.009353.001 GHz	2012-12-06	2	1.107	26.541
$\Box \oplus \leftrightarrow \sim$	2011.0.00397.S	J061200.23-062209.6	06:12:00.230	-06:22:09.600	7	0.5346	337.005352.989 GHz	2012-12-20	3	1.183	26.541
$\Box \oplus \leftrightarrow \sim$	2011.0.00397.S	J063027.81-212058.6	06:30:27.810	-21:20:58.600	7	0.5346	337.007352.992 GHz	2012-12-20	3	1.183	26.541
$\Box \Leftrightarrow \leftrightarrow \sim$	2011.0.00397.S	J035448.24-330827.2	03:54:48.240	-33:08:27.200	7	0.4848	337.026353.011 GHz	2012-12-20	3	1.128	26.541

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Position	🖓 Energy	🖓 Project	E Publication	© Observation		Redshift 0	estimated -
Source name	Frequency	Project code	BibCode	Observation Date		9	10
ALMA source name	Band	Project Title	Publication Title	Polarisation Type	CO v=0 5-4 H13CN v=0 J=6- CI 3P1-3P0	CO v=0 6-5 H2O v2=1 1(1,0)-1(0,1) HCI J=1-0,F1=5/2-3/2	HDO 1(1,1)-0(0,0) CH3OH vt=1 18(4,15)- CO v=0 7-6 H2O v=0 2(1,1)-2(0,2)
RA Dec	Spectral resolution	Project abstract	Abstract	Member ous id	6 5	,0)-1(0,1) ;5/2-3/2	.0(0,0) L 18(4,15)-19(L 18(1,15)-19(L,1)-2(0,2)
Galactic	Continuum sensitivity	PI Full Name	First Author	Object type	60	0 GHz	800 GHz
Target List	Line sensitivity (10 km/s)	Proposal authors	Authors	Public data only		9	10
Angular Resolution		Science keyword		Calibration observations			
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Max. Recoverable Scale			allowed operat	tors: =, <, >,, *, ?,	ublications	Ang. res.	Min. vel. res.
	1.S Fomalhaut b	22:57:38.685 -29:37:12.616 7	0.1181 34	13.077358.839 GHz 2012-12-06	2	arcsec ▼ 1.047	km/s ▼ 0.816
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		03:54:48.240 -33:08:27.200 7		37.026353.011 GHz 2012-12-20	3	1.128	26.541

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Spectral resolution	Project abstract	Abstract	Member ous id	2(1,2)	=3-2	-3 -2
Galactic	PI Full Name	First Author	Object type	00 GHz	250 GHz	300 GHz
Line sensitivity (10 km/s)	Proposal authors	Authors	 Public data only Calibration observations 	~	9	10
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↔ ~ 🖾 (2019.2.00112.S) Orion-KL	05:35:14.160 -05:22:21.500 6	0.2564	257.962261.75 GHz 2021-04-28	0	arcsec ▼ 4.562	km/s ▼ 0.561
↔ ~ 2019.2.00112.S Orion-KL	05:35:14.160 -05:22:21.500 6	0.2220	237.962241.75 GHz 2021-04-30	0	4.995	0.608
$\oplus \leftrightarrow \sim$ (2019.2.00094.S) ORS-9	05:35:14.600 -05:18:47.000 7	0.9623	277.862289.905 GHz 2021-04-22	0	4.194	0.076
	05:35:25.400 -05:24:33.000 7	0.9623	277.862289.905 GHz 2021-04-22	0	4.201	0.076
$\Leftrightarrow \leftrightarrow \sim \boxtimes$ (2019.2.00094.S) ORS-4	05:35:16.300 -05:20:43.000 7	0.9623	277.862289.905 GHz 2021-04-22	0	4.203	0.076
$\leftrightarrow \leftrightarrow \sim \bowtie$ (2019.2.00094.S) ORS-7	05:35:22.200 -05:25:10.000 7	0.9623	277.862289.905 GHz 2021-04-22	0	4.201	0.076

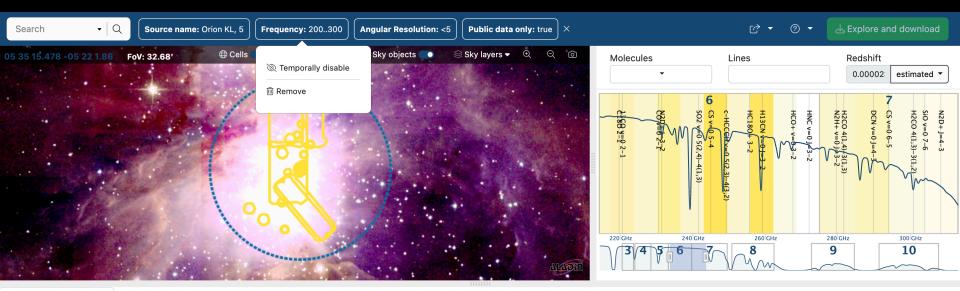


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© Observations (79)	○ Submillimetre ▼ SPIRE-color ▼ native ▼ <u>Remove</u>				
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			57.962261.75 GHz 2021-04-	28 0	4.562 0.561
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	05:35:14.600 -05:18:47.000 7	0.9623 277	77.862289.905 GHz 2021-04-	22 0	4.194 0.076
$\bigcirc \Leftrightarrow \leftrightarrow \sim \boxtimes \tag{2019.2.00094.S} \text{ ORS-5}$	05:35:25.400 -05:24:33.000 7	0.9623	77.862289.905 GHz 2021-04-	22 0	4.201 0.076
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Queried observations can be further filtered below or temporarily disabled above

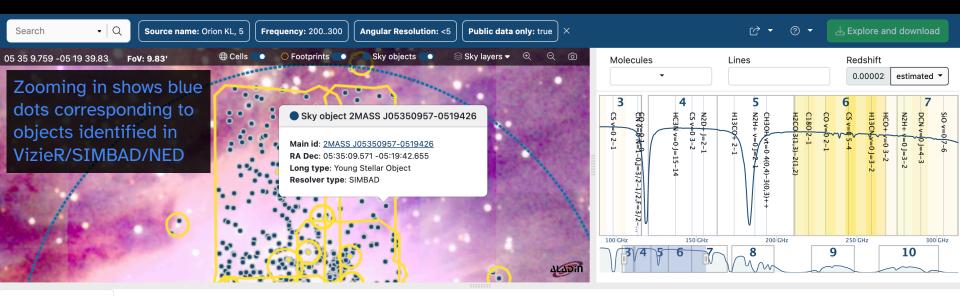
□ €	\rightarrow \leftrightarrow		↓ Project code	ALMA source name	RA h:m:s ▼	Dec d:m:s ▼	Band	Cont. sens. 	Frequency support	Release date	Publications	Ang. res.	Min. vel. res. km/s ▼
_ €	→ ↔	~	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2564	257.962261.75 GHz	2021-04-28	0	4.562	0.561
□ €	\rightarrow \leftrightarrow	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2220	237.962241.75 GHz	2021-04-30	0	4.995	0.608
_ €	\rightarrow \leftrightarrow	\sim	2019.2.00094.S	ORS-9	05:35:14.600	-05:18:47.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.194	0.076
_ €	\rightarrow \leftrightarrow	\sim	2019.2.00094.S	ORS-5	05:35:25.400	-05:24:33.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076
_ €	\rightarrow \leftrightarrow	\sim	2019.2.00094.S	ORS-4	05:35:16.300	-05:20:43.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.203	0.076
□ €	\rightarrow \leftrightarrow	\sim	2019.2.00094.S	ORS-7	05:35:22.200	-05:25:10.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076



Observations (52)

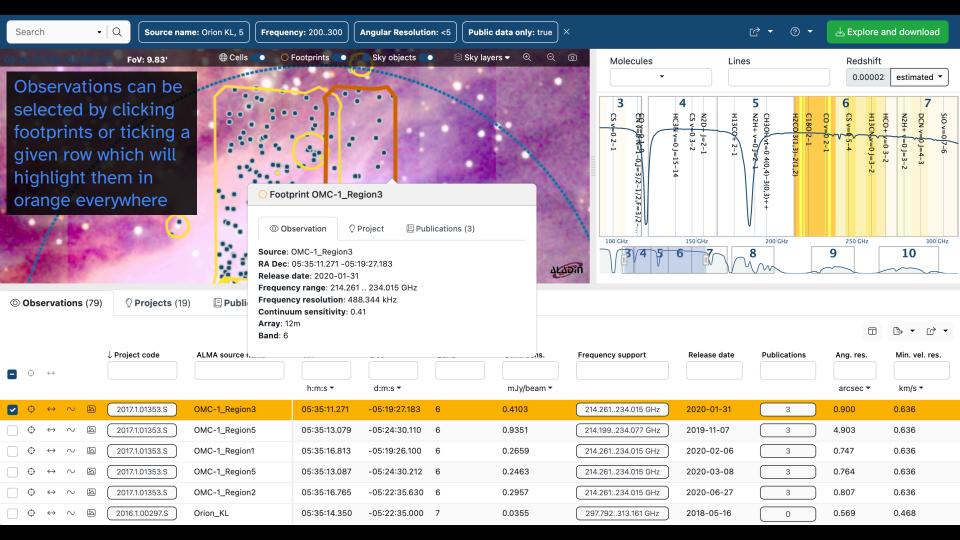
<u>Auteried</u> observations can be further filtered below or temporarily disabled above

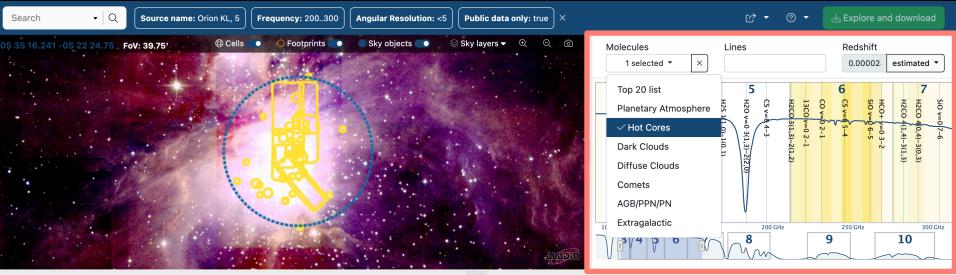
•	\leftrightarrow		↓ Project code	ALMA source name	RA h:m:s ▼	Dec d:m:s ▼	Band	Cont. sens. 	Frequency support	Release date	Publications	Ang. res. arcsec ▼	Min. vel. res. <1 ✔ × km/s ▼
•	\leftrightarrow	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2564	257.962261.75 GHz	2021-04-28	0	4.562	0.561
•	\leftrightarrow	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2220	237.962241.75 GHz	2021-04-30	0	4.995	0.608
•	\leftrightarrow	\sim	2019.2.00094.S	ORS-9	05:35:14.600	-05:18:47.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.194	0.076
Φ	\leftrightarrow	\sim	2019.2.00094.S	ORS-5	05:35:25.400	-05:24:33.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076
•	\leftrightarrow	\sim	2019.2.00094.S	ORS-4	05:35:16.300	-05:20:43.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.203	0.076
•	\leftrightarrow	\sim	2019.2.00094.S	ORS-7	05:35:22.200	-05:25:10.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076



 \bigcirc **Projects** (19) \square **Publications** (60)

•	\leftrightarrow		↓ Project code	ALMA source name	RA	Dec	Band	Cont. sens.	Frequency support	Release date	Publications	Ang. res.	Min. vel. res.
					h:m:s ▼	d:m:s ▼		mJy/beam ▼				arcsec 💌	km/s ▼
•	\leftrightarrow	\sim \(\Delta\)	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2564	257.962261.75 GHz	2021-04-28	0	4.562	0.561
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Φ	\leftrightarrow	\sim \begin{array}{c} \begin{array}{c} \end{array}	2019.2.00094.S	ORS-9	05:35:14.600	-05:18:47.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.194	0.076
Φ	\leftrightarrow	\sim \(\beta\)	2019.2.00094.S	ORS-5	05:35:25.400	-05:24:33.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076
•	\leftrightarrow	\sim \begin{array}{c} \begin{array}{c} \end{array}	2019.2.00094.S	ORS-4	05:35:16.300	-05:20:43.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.203	0.076
Φ	\leftrightarrow	\sim \	2019.2.00094.S	ORS-7	05:35:22.200	-05:25:10.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076

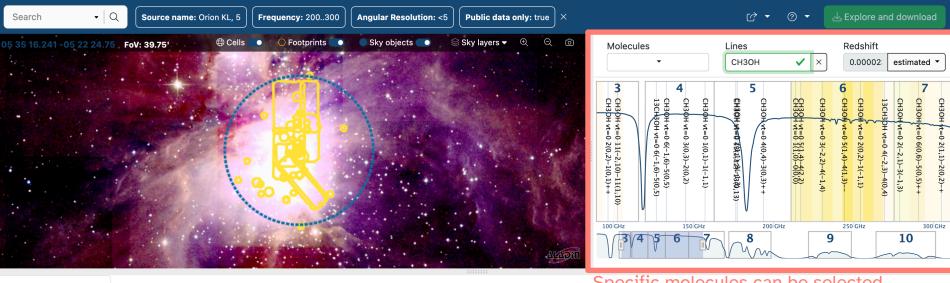




 \bigcirc **Projects** (19) \square **Publications** (60)

Molecules can be selected from a category to be shown in the spectral overview

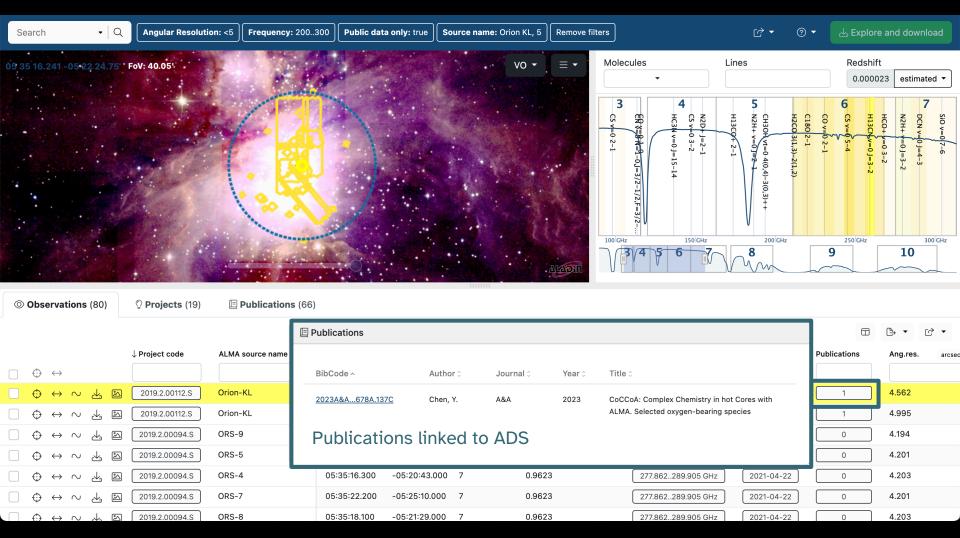
	\rightarrow \leftrightarrow		↓ Project code	ALMA source name	RA h:m:s ▼	Dec d:m:s •	Band	Cont. sens. mJy/beam ▼	Frequency support	Release date	Publications	Ang. res.	Min. vel. res. km/s ▼
_ €	\leftrightarrow	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2564	257.962261.75 GHz	2021-04-28	0	4.562	0.561
_ €	$\leftrightarrow \leftrightarrow$	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2220	237.962241.75 GHz	2021-04-30	0	4.995	0.608
_ €	\leftrightarrow	\sim	2019.2.00094.S	ORS-9	05:35:14.600	-05:18:47.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.194	0.076
_ €	$\leftrightarrow \leftrightarrow$	\sim	2019.2.00094.S	ORS-5	05:35:25.400	-05:24:33.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076
	\leftrightarrow	\sim	2019.2.00094.S	ORS-4	05:35:16.300	-05:20:43.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.203	0.076
	\leftrightarrow	\sim	2019.2.00094.S	ORS-7	05:35:22.200	-05:25:10.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076



 \bigcirc **Projects** (19) \square **Publications** (60)

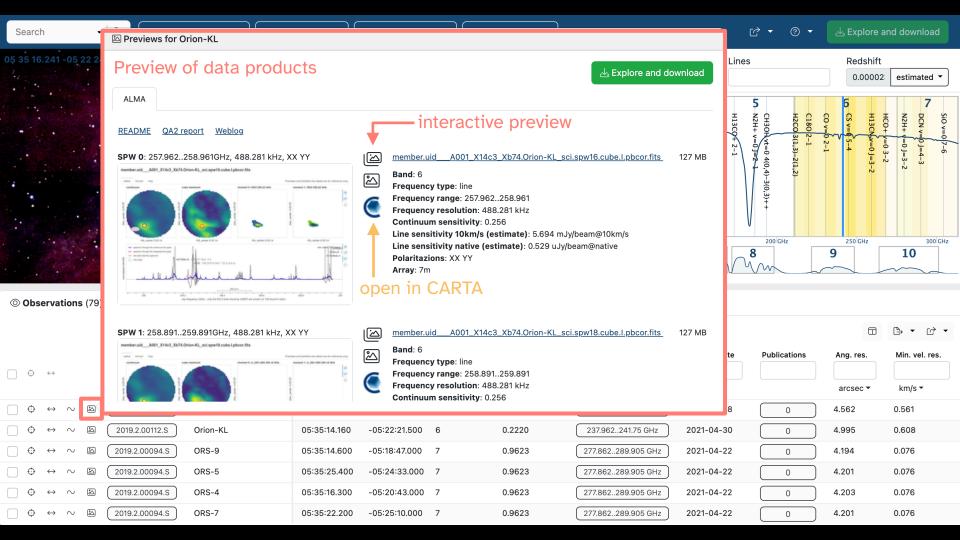
Specific molecules can be selected to be shown in the spectral overview

•	\leftrightarrow	↓ Project code	ALMA source name	RA	Dec	Band	Cont. sens.	Frequency support	Release date	Publications	Ang. res.	Min. vel. res.
				h:m:s ▼	d:m:s ▼		mJy/beam ▼				arcsec 💌	km/s ▼
Φ	\leftrightarrow \sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2564	257.962261.75 GHz	2021-04-28	0	4.562	0.561
Φ	\leftrightarrow \sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2220	237.962241.75 GHz	2021-04-30	0	4.995	0.608
•	\leftrightarrow \sim	2019.2.00094.S	ORS-9	05:35:14.600	-05:18:47.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.194	0.076
•	\leftrightarrow \sim	2019.2.00094.S	ORS-5	05:35:25.400	-05:24:33.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076
•	\leftrightarrow \sim	2019.2.00094.S	ORS-4	05:35:16.300	-05:20:43.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.203	0.076
Φ	\leftrightarrow \sim	2019.2.00094.S	ORS-7	05:35:22.200	-05:25:10.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076



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	Projects Public	cations			3 G	4 N2D HC3	H2C0 H2C0 -5 N2H+ H13C0	6 6 6	7 DCN 1 N2H+
	Project code 0	Title 0	Abstract 0		/=0 2-:	0+ J=2-1 7=0 3-2 7=0 J=	203(1,3)- 301 vt=0 1+ v=0 J=		
© Observations	<u>2019.2.00112.S</u>	Moving Past Small Number Statistics in Astrochemistry: An ACA Molecular Survey of 25 Hot Cores	Studies of chemical evolution have tradition small number of exceptionally molecularly ri due to the historical difficulty in detecting c molecules. This small sample size biases ou 'standard' chemical evolution, and prevents chemical models to standard conditions. Sir have attempted to address this issue by sur larger sample sizes have suffered from extrr due to the very small angular size of typical cores. Here, we propose to exploit both the resolution of ALMA to conduct a distance-li five hot core sources. We target several spe designed to provide the maximal scientific r	ich and bright sources omplex interstellar in understanding of calibration of ngle-dish surveys that veying substantially eme beam dilution, chemically rich hot sensitivity and spatial mited survey of twenty ectral windows eturn, and will use the		1 =15-14 1-0_J=3/2-1/2,F=3/2	1 1 1 1 1 1 1 1 1 1 1 1 1 1	250 CHz	0 3-2 0 3-2 0 0 0 0 0 0 0 0 0 0 0 0 0
	Click to fin	d	results to calibrate several of the industry-s identify critical areas in which these models			F	Delesse dete	Dublications	÷ ۲ ۲ •
$\Box \oplus \leftrightarrow$	similar pro	jects	These ACA observations complement an ap array for an identical set of sources and spe	· · · ·	i. mJy/beam ▼	Frequency support	Release date	Publications	Ang.res. arcsec
$\square \oplus \leftrightarrow \sim$				-		257.962261.75 GHz	2021-04-28		4.562
$\bigcirc \leftrightarrow \sim$	<u>ک</u> 2019.2.00112	2.S Orion-KL	05:35:14.160 -05:22:21.500	6 0.2220		237.962241.75 GHz	2021-04-30		4.995
$\Box \Leftrightarrow \leftrightarrow \sim$	كم الأ	4.S ORS-9	05:35:14.600 -05:18:47.000	7 0.9623		277.862289.905 GHz	2021-04-22	0	4.194
$\Box \oplus \ \leftrightarrow \ \sim$	よ。 「2019.2.0009	4.S ORS-5	05:35:25.400 -05:24:33.000	7 0.9623		277.862289.905 GHz	2021-04-22	0	4.201
$\Box \Leftrightarrow \leftrightarrow \sim$	كم الم 2019.2.0009	4.S ORS-4	05:35:16.300 -05:20:43.000	7 0.9623		277.862289.905 GHz	2021-04-22	0	4.203
$\Box \Leftrightarrow \leftrightarrow \sim$	丛, 釣 2019.2.0009	4.S ORS-7	05:35:22.200 -05:25:10.000	7 0.9623		277.862289.905 GHz	2021-04-22	0	4.201
$\bigcirc \oplus \leftrightarrow \sim$	上 图 2019.2.0009	4.S ORS-8	05:35:18.100 -05:21:29.000	7 0.9623		277.862289.905 GHz	2021-04-22	0	4.203

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	2	} •••••		3350 7(6)-6(5), CH2CHEN V11- H13CN V=0 J=3 NH2CHO 13(2,1 SO2 V=0 32(4,2 SO 35/gmax=0 HC155V=0 3-2	HCQ 3(0,3 G g-CH3CH2 SO2 v=0 3	, <u> </u>	13С430Н 13С43СН NH2CH0
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↓ Project code ALMA	source name RA h:m:s -	Dec d:m:s - Bar	nd Cont.sens. mJy/beam -	Frequency support	Release date	Publications A	Ang.res. arcsec
\checkmark ↔ \sim \checkmark \succeq 2019.2.00112.S Orion-	KL 05:35:14.160	0 -05:22:21.500 6	0.2564	257.962261.75 GHz	2021-04-28	1 4	1.562
$\bigcirc \leftrightarrow \sim \swarrow \boxtimes 2019.2.00112.s \qquad Orion-$	KL 05:35:14.160	0 -05:22:21.500 6	0.2220	237.962241.75 GHz	2021-04-30	1 4	1.995
\bigcirc \leftrightarrow \sim \swarrow \bigtriangleup 2019.2.00094.S ORS-S	05:35:14.60	0 -05:18:47.000 7	0.9623	277.862289.905 GHz	2021-04-22	0 4	1.194
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□ ↔ ↔ ∼ ৬ [2019.2.00094.S] ORS-4	05:35:16.30	0 -05:20:43.000 7	0.9623	277.862289.905 GHz	2021-04-22	0 4	1.203
↔ ∼ , , ∑ [2019.2.00094.S] ORS-7	05:35:22.20	0 -05:25:10.000 7	0.9623	277.862289.905 GHz	2021-04-22	0 4	1.201
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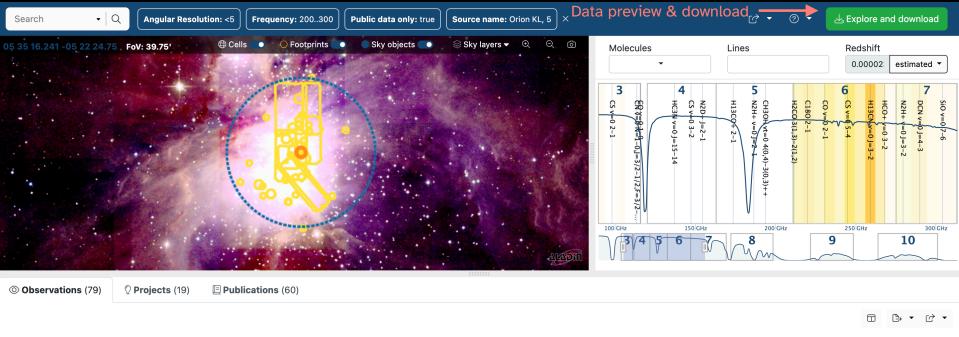


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interactive preview example

Previews and tentative line labels are for reference only. native binned help moment 0: HDO 258.22 GHz moment 1: HDO 258.22 GHz cube maximum continuum ÷ 09 Dec_center -5:22:22 Dec_center -5:22:22 Dec_center -5:22:22 Dec_center -5:22:22 0 \square \bigcirc RA center 5:35:14 RA center 5:35:14 RA center 5:35:14 RA_center 5:35:14 spectrum through the continuum-flux peak visr used: 5,00-km _ spectrum through the integrated-flux peak _ SO2v=0 2 0 full-cube total-flux spectrum _ H13CNv2=1 99 normalised intensity rms noise HC15Nv=0 258157 GHz 3-2 HDO 258.22376 GHz 7(5,3)-8(4,4) Ð 200 km/s 258 258.2 258.4 258.6 258.8 259

sky frequency (GHz) - only the first 5 lines found by ADMIT are shown (of 139 found in total) -



				\downarrow Project code	ALMA source name	RA	Dec	Band	Cont. sens.	Frequency support	Release date	Publications	Ang. res.	Min. vel. res.
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						h:m:s ▼	d:m:s ▼		mJy/beam ▼				arcsec 🔻	km/s ▼
	Φ	\leftrightarrow	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2564	257.962261.75 GHz	2021-04-28	0	4.562	0.561
	Φ	\leftrightarrow	\sim	2019.2.00112.S	Orion-KL	05:35:14.160	-05:22:21.500	6	0.2220	237.962241.75 GHz	2021-04-30	0	4.995	0.608
	Φ	\leftrightarrow	\sim	2019.2.00094.S	ORS-9	05:35:14.600	-05:18:47.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.194	0.076
	Φ	\leftrightarrow	\sim	2019.2.00094.S	ORS-4	05:35:16.300	-05:20:43.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.203	0.076
	Φ	\leftrightarrow	\sim	2019.2.00094.S	ORS-5	05:35:25.400	-05:24:33.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076
	Φ	\leftrightarrow	\sim	2019.2.00094.S	ORS-7	05:35:22.200	-05:25:10.000	7	0.9623	277.862289.905 GHz	2021-04-22	0	4.201	0.076

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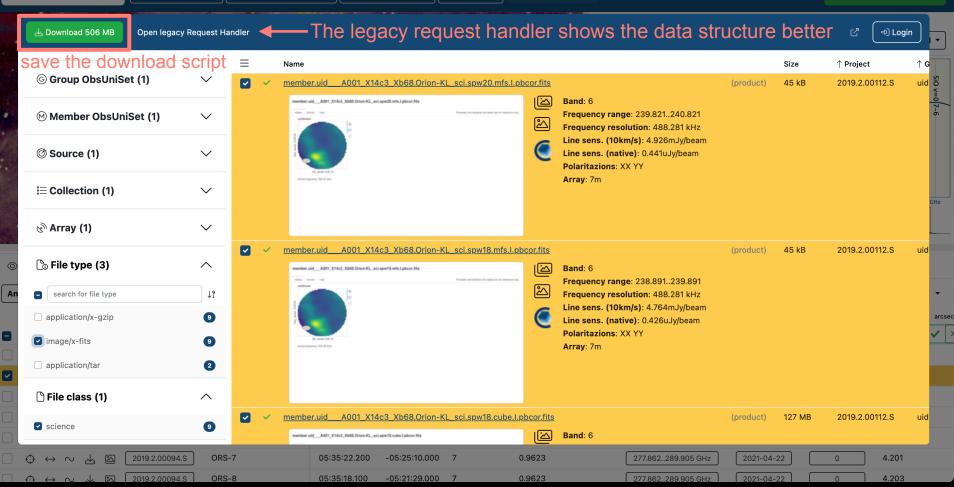
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Frequency: 200..300 Source name: Orion KL, 5 Public data only: true Remove filters

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ALMA Request Handler

Anonymous User: Request #2166926188317 🗹

Legacy download interface

Request Title: <u>click to edit</u>
Download Selected

Need to select readme+auxiliary+raw to

- be able to run pipeline calibration/imaging

🗹 readme 🗹 product 🗹 auxiliary 🗹 raw 🗹 raw (semipass) 🗹 external

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ALMA Request Handler

Anonymous User: Request #1658068127545 ***** Request Title: <u>click to edit</u>

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ALMA Data Products

+ TARAO STORAGE AND A STADIA

- The downloaded and unpacked data fall into a standardized directory structure project_id/
- Note: depending on the cycle, the delivered products may vary...

Check for each cycle: <u>https://almascience.org/processing/qa2-data-products</u>

ALMA Data Products

• Possible FITS images and what they are

spw##.cube.I.
spw##.mfs.I.
spw##_##_##_##.cont.I.
spw##_##_##_##.cont.I.alpha.
spw##_##_##_##.cont.I.tt0.

spw##_##_##.cont.I.tt1.

spw##_##_##.cont.IQUV.

A spectral image cube of a single spectral window A continuum image for a single spectral window An aggregate bandwidth or continuum image A spectral index image An image containing the zeroth Taylor term for a continuum image An image containing the first Taylor term for a continuum image An aggregate bandwidth or continuum full Stokes cube



ALMA Regional Centre || Allegro

ALMA Data Products

T TARA A TOTAL AND A TANK

- Possible FITS images and what they are
- *.mask.fits
 *pb.fits or *.flux.fits
 *pbcor.fits
 *sd.im.fits
 .mfs.A. or *mfs.POLA*
 .mfs.P. or *mfs.POLI*
- The mask that was used when the image was created The primary beam response for a field A primary-beam corrected image A single dish image A polarization angle map A linear polarization intensity map



llearo

Check data quality

T. THERE AND AN AND AND AN A READAN

- Assess data quality through the weblog created by the pipeline
- The weblog resides in the qa folder as a tar file that must be unpacked

```
hember_ouss_id/
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```
README
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```
product/
```

```
calibration/
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```
*.weblog.tgz
```

```
script/
```

```
log/
raw/
```

Go into the qa folder
unpack the *.weblog.tgz file
> tar -xvzf *.weblog.tgz
Go inside the newly-created
pipeline-timestamp folder
Go inside the html folder
open the weblog with Firefox
> firefox index.html



 Note: Sometimes Firefox has issues opening the weblog. Alternatively one can open the weblog from within CASA via h_weblog() <u>https://help.almascience.org/kb/articles/what-is-the-best-way-to-view-the-weblog</u>

Calibrated Visibilities

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- To retrieve calibrated visibilities for your own or archival data:
 - Request them via a Helpdesk ticket: <u>https://almascience.eso.org/local-news/requesting-calibrated-measurement-</u> <u>sets-in-europe</u>
 - If ARI-L products exist, the calibrated .MS files can be requested from the ARI-L team for at least 3 years after June 2022 (<u>https://sites.google.com/inaf.it/ari-l</u>)
 - o Recreate them yourself!

More information:

https://help.almascience.org/kb/articles/how-do-i-obtain-a-file-of-calibrated-visibilities-measurement-set-for-alma-data



Recreating Calibrated Visibilities

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• The calibration of both pipeline calibrated and manually calibrated ALMA data is performed by running *.scriptForPI.py in the script directory.

pipeline-calibrated MOUS

script/

- scriptForPI.py
- casa_pipescript.py
- casa_piperestorescript.py
- PPR*.xml or *pprequest.xml

manually-calibrated MOUS script/

scriptForPI.py

- The raw data in ALMA Science Data Model (ASDM) Format

- The correct version of CASA with the ALMA pipeline included

- scriptForCalibration.py (per EB)
 - scriptForFluxCalibration.py
- scriptForImagingPrep.py (sometimes)

YOU NEED:

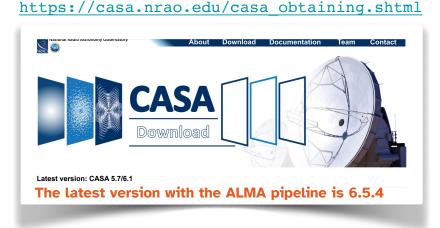


Obtain CASA

- Determine the version needed:
 - Cycle 5+: Find this information in the QA2 pdf report (in the qa directory)
 - At the top of the CASA log files (in the log directory or in qa/pipeline*/html)

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- If pipeline-calibrated: on the first page of the WebLog
- If manually-calibrated/early cycles:
 - in the **README** file
 - Near the top of the script scriptForCalibration.py
- Install the version with the ALMA Pipeline!





Run scriptForPI

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pipeline calibration

① Go to the script folder & launch CASA with the pipeline extension:

- > casa --pipeline
- @ execute the *.scriptForPI.py file:

CASA <x>: execfile('member.uid*.scriptForPI.py')

manual calibration

1 Go to the script folder & launch CASA <u>without</u> the pipeline extension:

- > casa
- @ execute the *.scriptForPI.py file:

CASA <x>: execfile('member.uid*.scriptForPI.py')



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After running scriptForPI

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- After successful completion of scriptForPI.py a new folder is created project_id/
 - └→ science goal ouss id/ └→ group ouss id/ └→ member ouss id/ -> README > product/ -> calibration/ → qa/ -> script/ -> log/ ▶ raw/ calibrated ↓ 1 calibrated .ms file per EB



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Typical Problems

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- "Not enough disk space"
 - run the script in space saving mode by starting it in the following way

> casa -c "SPACESAVING -N; execfile('scriptForPI.py')"

- N = 0 -> no space saving mode
- N = 1 -> all intermediate measurement sets *ms.split are deleted
- N = 2 -> all intermediate measurement sets *ms.split and *.ms are deleted
- N >= 3 -> all intermediate measurement sets *ms.split, *.ms, and *ms.split.cal are deleted
- "TypeError" or "NameError"
 - o most likely using the wrong CASA version



Typical Problems

- "No such file or directory"
 - change file extension names in the calibration folder from *.tar.gz to *.tgz

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- "The calibrated directory already exists"
 - This folder is created every time you execute the scriptForPI.py script, so remove it before you want to run the script again





Querying the archive programmatically

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- pyVO ADQL queries through Table Access Protocol (TAP) service
 - o <u>Documentation</u>
 - o Example notebooks for ALMA
- Astroquery
 - o <u>Documentation</u>
- ALminer: ALMA Archive Mining & Visualization Toolkit
 - o <u>Documentation</u>

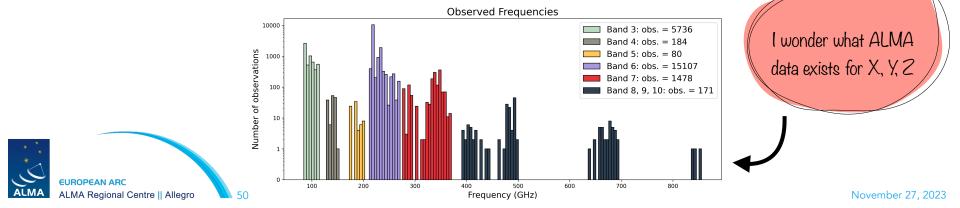


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Python-based code to effectively **query**, **analyse**, and **visualise** the ALMA Science Archive + **download** raw/products data



Where to find ALminer

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Documentation: https://alminer.readthedocs.io/

GitHub: https://github.com/emerge-erc/ALminer

You Tube I-TRAIN video: https://bit.ly/ALminer I-TRAIN video



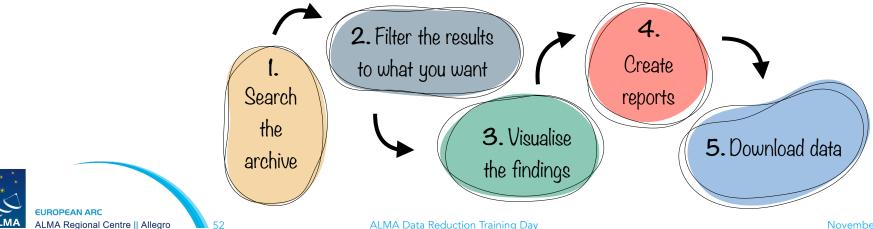


ALminer's extensive tutorial

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Static version at https://alminer.readthedocs.io/

Live Jupyter Notebook Slaunch Jupyter Notebook



References

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- ALMA Science Archive: <u>http://almascience.org/aq/</u>
- Science Archive Manual: <u>https://almascience.eso.org/documents-and-tools/</u> <u>latest/science-archive-manual</u>
- Science Archive Primer: <u>https://almascience.eso.org/documents-and-tools/</u> <u>cycle9/archive-primer</u>
- Data processing overview: <u>https://almascience.eso.org/processing</u>
- Petry et al. 2020: <u>https://ui.adsabs.harvard.edu/abs/2020Msngr.181...16P/abstract</u>



Next up:

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- Calibration fundamentals
- Imaging techniques
- Analysis & visualisation Tools

