Common Astronomy Software Applications

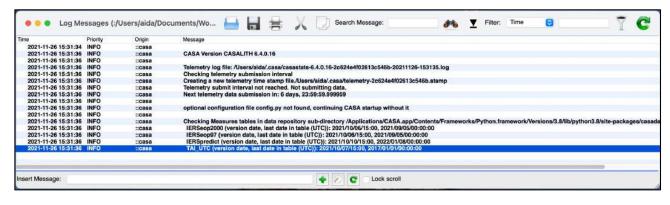
- Can process data from both single-dish and aperture-synthesis telescopes
- Primary data processing (calibration & imaging) software for ALMA & VLA
- Python based
 - Versions < 6 use Python 2.7
 - Newer versions use Python 3
- Available for Linux (RedHat) and Mac OS
 - → Website https://casa.nrao.edu/
 - → Guides https://casaguides.nrao.edu/
 - → Documentation
 - Versions 6.1 & earlier: https://casa.nrao.edu/casadocs
 - Versions 6.2 & later: https://casadocs.readthedocs.io/en/stable/



Starting CASA

After installation, to open CASA simply type casa in the terminal if you set up an alias. Otherwise type the full path.

→ Starting CASA will open a logger (and a log file):





casa --nologger
if you do not need
the logger GUI

→ And the terminal prompt:

```
optional configuration file config.py not found, continuing CASA startup without it

IPython 7.15.0 -- An enhanced Interactive Python.

Using matplotlib backend: MacOSX
Telemetry initialized. Telemetry will send anonymized usage statistics to NRAO.
You can disable telemetry by adding the following line to the config.py file in your rcdir (e.g. -/.casa/config.py):
telemetry_enabled = False
--> CrashReporter initialized.
CASA 6.4.0.16 -- Common Astronomy Software Applications [6.4.0.16]

CASA <1>: ||
```

CASA Basics

CASA Tasks

Tasks are executed to perform a single job (e.g. loading, plotting, flagging, calibrating) Each task contains a set of user-defined parameters

List of available tasks

taskhelp -> A more exhaustive list of tasks with descriptions
tasklist() -> Get an overview of available tasks, organized by category (removed in CASA 6)
+ More information about the tasks:

https://casadocs.readthedocs.io/en/stable/api/casatasks.html

Getting help on a task

inp <taskname> to get an overview of a given task and its input parameters

help <taskname> to get a detailed description of a given task and its input parameters

(use arrow keys to continue, press q to exit)

Based on: https://casaquides.nrao.edu/index.php?title=Getting_Started_in_CASA

CASA Basics

Executing a task

Interactively:

You may also do:

default (<taskname>) -> to set the parameters of a task to their default values
set individual parameters using a Python cparameter>=<value> syntax

Note: you can also simply set parameters without the default or tget steps but beware that you would be setting parameters globally!

Programmatically:

```
taskname(parameter1='', parameter2='', ...)
```

Based on: https://casaquides.nrao.edu/index.php?title=Getting_Started_in_CASA

CASA Basics

Parameters

grey: parameter has sub-parameters

green: sub-parameters

red: invalid value

blue: parameter altered from its default

Data selection syntax

spw='0:5~30;40~55,1:10~25;45~58,2'

Running scripts

In CASA: execfile('script_name.py')

In the terminal: casa -c script_name.py

vis	= 'data/sis14_twhy	a_calibrated_flagged.ms'
	and the same of th	# Name of input visibility file(s)
selectdata	# True	# Enable data selection parameters
field	= 11	# field(s) to select
Spw	= 11	# spw(s)/channels to select
timerange	= **	# Range of time to select from data
uvrange	= **	# Select data within uvrange
antenna	= 4.1	# Select data based on antenna/baseline
scan	= 1.1	# Scan number range
observation	= 11	# Observation ID range
intent	= 11	# Scan Intent(s)
datacolumn	= 11	# Data column to image(data,corrected)
imagename	= 11	# Pre-name of output images
imsize	= [100]	# Number of pixels
cell .	= []	# Cell size
phasecenter	= * * *	# Phase center of the image
stokes	= 'I'	# Stokes Planes to make
projection	= 'SIN'	# Coordinate projection
startmodel	= ''	# Name of starting model image
specmode	= 'mfs'	# Spectral definition mode (mfs,cube,cubedata, cubesource)
reffreq	= 11	# Reference frequency
gridder	= 'standard'	# Gridding options (standard, wproject, widefield, mosaic, awproject)
vptable	E 11	# Name of Voltage Pattern table
polimit	= 0.2	# PB gain level at which to cut off normalizations
deconvolver	= 'hogbom'	# Minor cycle algorithm (hogbom,clark,multiscale,mtmfs,mem,clarkstokes)
restoration	= True	# Do restoration steps (or not)
restoringbeam	= []	# Restoring beam shape to use. Default is the PSF main lobe
pocor	= False	# Apply PB correction on the output restored image
outlierfile	= ''	# Name of outlier-field image definitions
weighting	= 'nat'	# Weighting scheme (natural,uniform,briggs, briggsabs[experimental], briggsbwtaper[experimental])
niter	= 0	# Maximum number of iterations
usemask	= 'user'	# Type of mask(s) for deconvolution: user, pb, or auto-multithresh
mask	= 11	# Mask (a list of image name(s) or region file(s) or region string(s))
pbmask	= 0.0	# primary beam mask
fastnoise	= True	# True: use the faster (old) noise calculation. False: use the new improved noise calculations
restart	= True	# True : Re-use existing images. False : Increment imagename
savemodel	= 'none'	# Options to save model visibilities (none, virtual, modelcolumn)
calcres	= True	# Calculate initial residual image
calcpsf	= True	# Calculate PSF
psfcutaff	= 0.35	# All pixels in the main lobe of the PSF above psfcutoff are used to fit a Gaussian beam (the Clean beam)
parallel	= False	# Run major cycles in parallel

CASA <39>: inp tclean

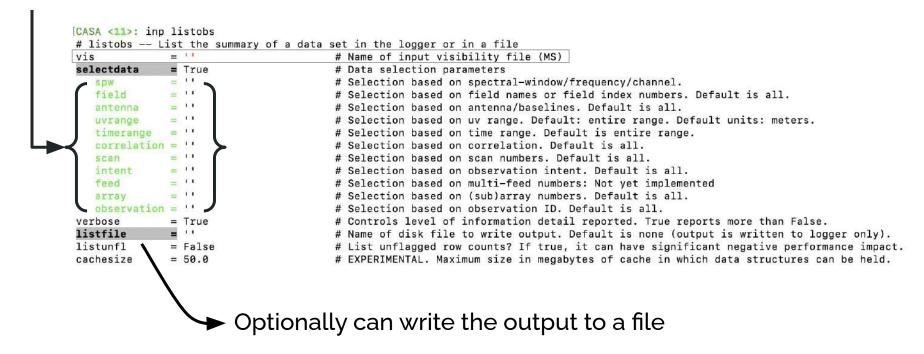
tolean - Radio Interferometric Image Reconstruction

Data Inspection with CASA

listobslist the contents of measurement setplotantsplot the location of antennasplotmsinspect/flag visibilities interactivelyimviewview/inspect images interactively

listobs: lists the contents of measurement set

Can select a subset of the measurement set



listobs: lists the contents of measurement set

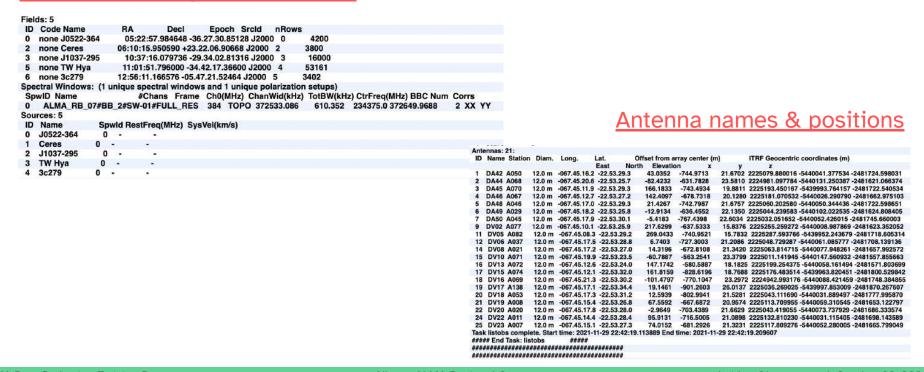
Example > listobs(vis='sis14_twhya_calibrated_flagged.ms')

sequence of observations

MeasurementSet Name: /Users/aida/Documents/Work/Leiden/Allegro/Events/202111_Data_Reduction_Day/data/sis14_twhya_calibrated_flagged.ms MS Version 2					
Observer: cqi Project: uid:///	4002	X327408/X6f			
servation: ALMA					
mputing scan and subscan p					
		time = 5647.68 sec			
Observed from 19-Nov-2012/0	17:36	57.0 to 19-Nov-20	7/2/09:11:04.7 (UTC)		
ObservationID = 0 ArrayID	-0				
ate Timerange (UTC)		n Fldld FieldName	nRows Spwids Average Interval(s) Scanintent		
9-Nov-2012/07:36:57.0 - 07:39:			4200 [0] [6.05] [CALIBRATE BANDPASS#ON_SOURCE, CALIBRATE_PHASE#ON_SOURCE, CALIBRATE_WVR#ON_SOURCE		
07:44:45.2 - 07:47:01.2	7.7	2 Ceres	3800 [0] [6.05] [CALIBRATE AMPLI#ON SOURCE, CALIBRATE PHASE#ON SOURCE, CALIBRATE WVR#ON SOURCE]		
07:52:42.0 - 07:53:47.6		3 J1037-295	1900 [0] [6.05] [CALIBRATE PHASE#ON SOURCE, CALIBRATE WVR#ON SOURCE]		
07:56:23.5 - 08:02:11.3	12	5 TW Hya	8514 [0] [6.05] [OBSERVE TARGET#ON SOURCE]		
08:04:36.3 - 08:05:41.9	14	3 J1037-295	1900 [0] [6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
08:08:09.6 - 08:13:57.3	16	5 TW Hya	10360 [0] [6.05] [OBSERVE_TARGET#ON_SOURCE]		
08:16:20.6 - 08:17:26.2	18	3 J1037-295	2100 [0] [6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
08:19:53.9 - 08:25:41.7	20	5 TW Hya	10321 [0] [6.05] [OBSERVE_TARGET#ON_SOURCE]		
08:28:17.1 - 08:29:22.6		3 J1037-295	2100 [0] [6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
08:32:00.5 - 08:37:48.2	24	5 TW Hya	10324 [0] [6.05] [OBSERVE_TARGET#ON_SOURCE]		
08:40:11.9 - 08:41:17.4		3 J1037-295	2100 [0] [6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
08:43:45.6 - 08:49:33.4		5 TW Hya	9462 [0] [6.05] [OBSERVE_TARGET#ON_SOURCE]		
08:51:57.1 - 08:53:02.6		3 J1037-295	1900 [0] [6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
08:58:12.0 - 09:00:28.1	170	6 3c279	3402 [0] [6.05] [CALIBRATE_BANDPASS#ON_SOURCE,CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
09:01:35.7 - 09:02:41.2		3 J1037-295	1900 [0] [6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]		
09:05:15.6 - 09:07:31.6	75.00	5 TW Hya 3 J1037-295	4180 [0] [6.05] [OBSERVE_TARGET#ON_SOURCE] 2100 [0] [6.05] [CALIBRATE PHASE#ON SOURCE, CALIBRATE WVR#ON SOURCE]		
09:09:59.1 - 09:11:04.7					

listobs: lists the contents of measurement set

<u>List of fields & spectral windows</u>



Getting started on the Allegro computers

```
Go to the analysis folder in the project directory
   > cd
   /allegro1/allegro/home/your username/open ALMA DRT2023/analysis/your username
   Make two folders
   > mkdir imaging
   > mkdir analysis tools
3. Copy data from the 'archive' folder to your own folder
   > cp -r
   ../../archive/DRT2023/TW hydra/sis14 twhya calibrated flagged.ms.contsub
   imaging/.
   > cp -r ../../archive/DRT2023/TW hydra/twhya n2hp.image analysis tools/.
   > cp -r ../../archive/DRT2023/TW hydra/sis14 twhya cont.image analysis tools/.
   > cp -r ../../archive/DRT2023/TW hydra/*.fits analysis tools/.
4. Copy scripts from the 'scripts' folder to your own folder
   > cp ../../scripts/Imaging*.py imaging/.
   > cp ../../scripts/analysis*.py analysis tools/.
  Go to the imaging folder and open CASA
   > cd imaging
   > nice +10 env -u PYTHONPATH -u LD LIBRARY PATH casapy-660
```